

Vrije Universiteit Brussel

IceCube AGN/GRB searches at the IIHE

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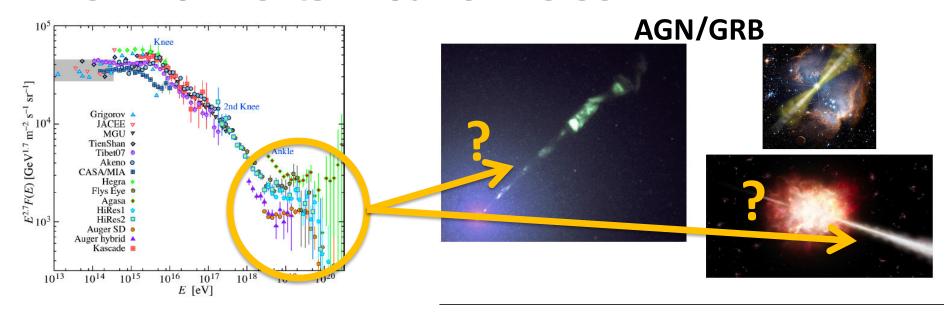






IceCube AGN/GRB searches: Why?

 What is going on at the most energetic environments in our Universe?

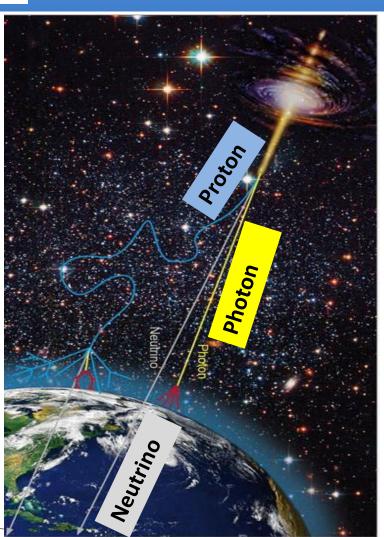


 Where do our Ultra-High Energy Cosmic Rays come from?

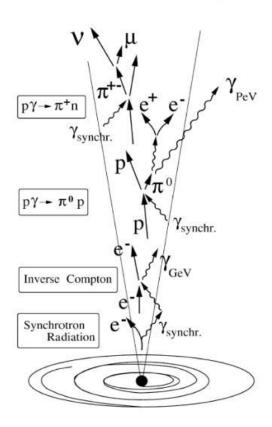




Probing AGN/GRB Why Neutrinos??



Processes in the jet

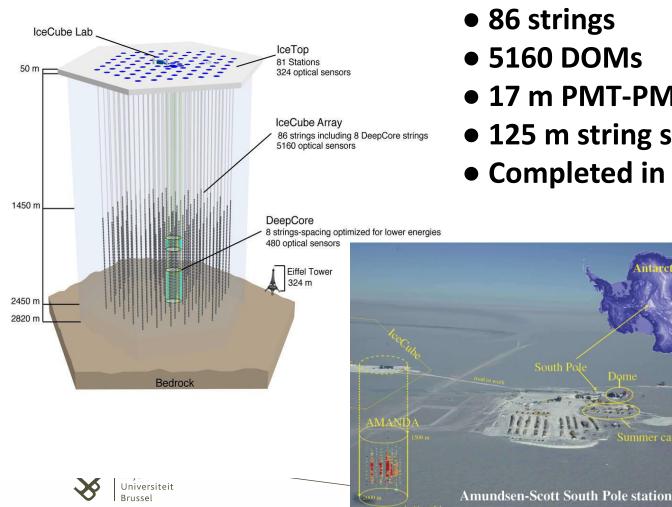


High-energy nuclei, γ and u





How to detect cosmic neutrinos? The IceCube detector



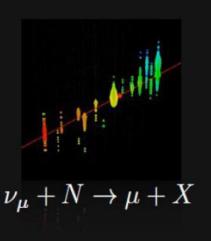
• 1 km³ volume

- 17 m PMT-PMT spacing per string
- 125 m string spacing
- Completed in Dec. 2010



Neutrino signatures in IceCube

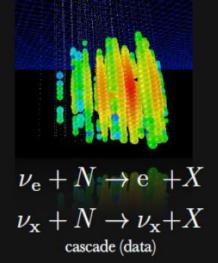
CC Muon Neutrino



track (data)

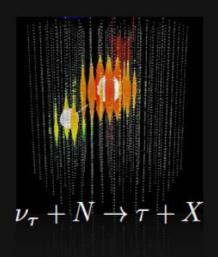
factor of ≈ 2 energy resolution < 1° angular resolution

Neutral Current /Electron Neutrino



≈ ±15% deposited energy resolution ≈ 10° angular resolution (at energies ≥ 100 TeV)

CC Tau Neutrino



"double-bang" and other signatures (simulation)

(not observed yet)

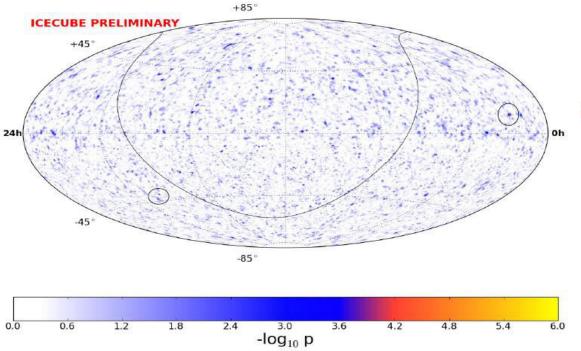
(C. Kopper)





IceCube Point Source Analysis





4 years of detector data, no evidence of point sources was found.

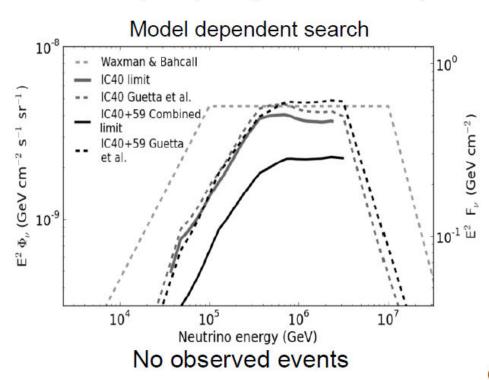
Most significant spots post trials P-values: Northern sky 22.6% Southern Sky 44.0%

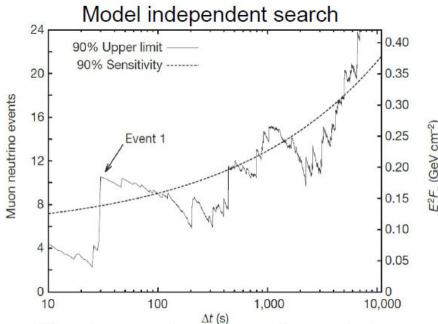




GRB searches in IceCube

Nature 484 (2012) 351 (IC-40/IC-59 data)





No observed excess of events in coincidence with GRBs (event at $\Delta t = 30$ s consistent with bckg cosmic ray event.)

Limits exclude all tested models with standard parameters.

Models are being revisited => Recalculations.

New publication with IC-79/IC-86 data in progress!



The Brussels IIHE AGN/GRB event selection

- No signal seen by IceCube so far.
- Up to now most pre-selections were based on <u>hard cuts</u> on certain parameters.
- We try to <u>retrieve</u> some of the possible signal events by assigning <u>weights</u> to the quality parameters of different track reconstruction algorithms.
- A first pre-selection cut is performed on a <u>linear sum of these</u> weights: Hence, <u>if one reconstruction is bad we do not</u> <u>immediately throw away the event.</u>





The Brussels IIHE AGN/GRB event selection

Example:

Log-likelihood quality parameter: RlogL

Causality parameter: Beta=v/c

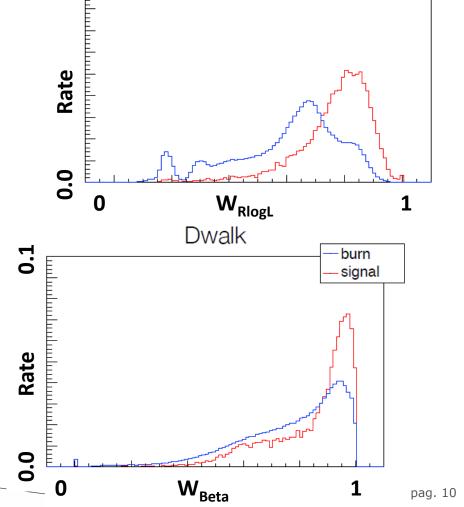
$$w_{R\log L} = \frac{6}{R\log L}$$

$$w_{Beta} = \begin{cases} \beta & (\beta < 1) \\ 2 - \beta & (\beta > 1) \end{cases}$$

2 hour burn sample in which **Burn:**

no signal is expected.

Signal: Nugen simulation



MPE



burn

signal



The Brussels IIHE AGN/GRB event selection

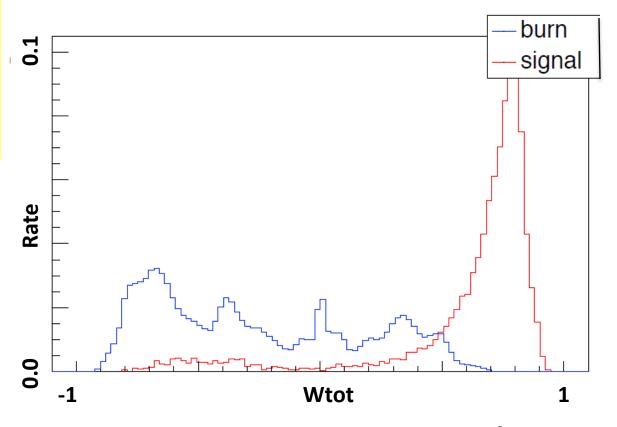
Example: Total weight distribution

Weight:
$$W_{Tot} = \sum_{reco,i} \pm a_i W_i$$

$$+: \theta > 90^{\circ}$$

$$-: \theta < 90^{\circ}$$

Cutting at 80% signal leaves only 1% background!!



Lionel Brayeur



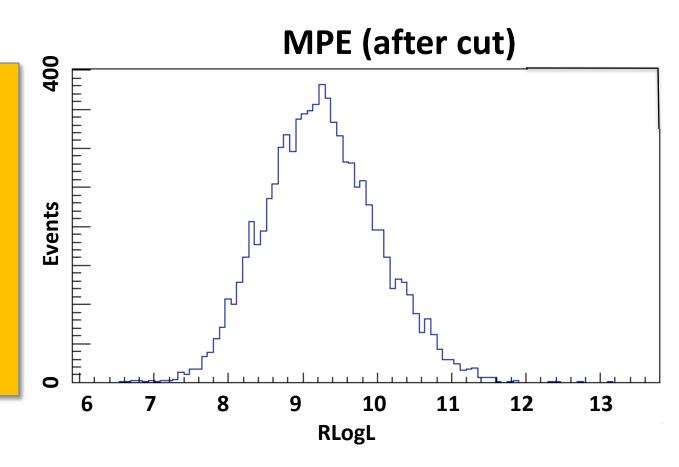


The Brussels IIHE AGN/GRB event selection

Many events which would not survive a typical IceCube pre-cut

RLogL < 8 - 10

are still maintained with a very good efficiency!!

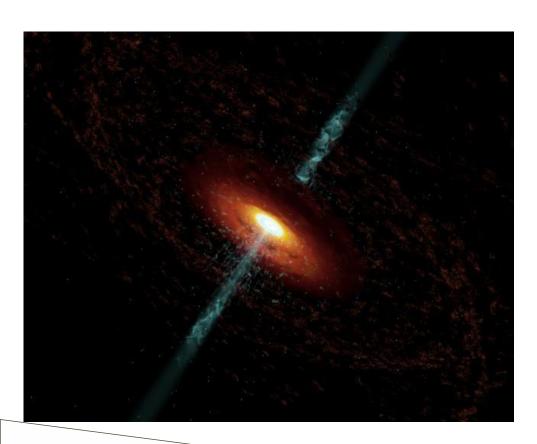




Lionel Brayeur



A specific type of AGN will be considered for the IIHE analysis



General Blazar properties:

- Very strongly polarized variable emission over a wide range of frequencies
- Neutrino production:

$$p_{jet}^{+} + \gamma \rightarrow \pi^{+} + n$$

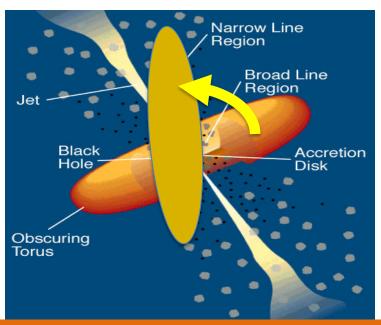
$$\pi^{+} \rightarrow \mu^{+} + \nu_{\mu}$$

$$\rightarrow e^{+} + \nu_{e} + \nu_{\mu} + \overline{\nu}_{\mu}$$





The Brussels selection: Search for a specific type of Blazar



- At a distance of a few parsec a dust torus is formed.
- This torus does not have to be perpendicular to the jet!
- Enhanced neutrino production can occur through the jet- torus interaction:

$$p_{jet}^{+} + p_{torus}^{+} \to X$$

$$X \to \pi^{+} \to \mu^{+} + \nu_{\mu}$$

$$\to e^{+} + \nu_{e} + \nu_{\mu} + \overline{\nu}_{\mu}$$

- Radio waves will pass through the torus without interaction.
- Emission at higher frequencies will be obscured



Search for AGN with high radio

luminosity and low luminosities at

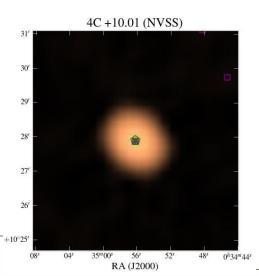
higher frequencies!!

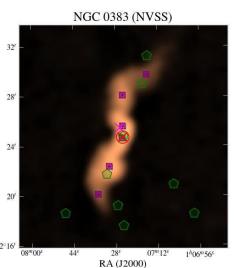


Nijmegen AGN radio catalogue

S. van Velzen, et al., Astronomy & Astrophysics 544 (2012) A18 "Radio Galaxies of the Local Universe", arxiv:1206.0031

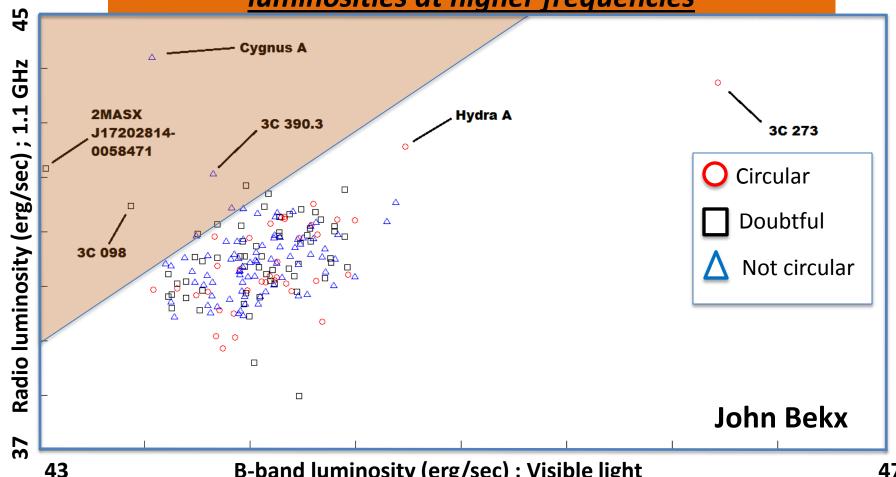
- Goal: Select AGN within 100 Mpc which could be responsible for UHECRs
- 407 entries in "Jets & Lobes" are investigated
- First categorization: Radio morphology







Search for AGN with High radio luminosity and low luminosities at higher frequencies



B-band luminosity (erg/sec) ; Visible light

47



Conclusion:

407 entries in "Jets & Lobes" are investigated
First categorization: Circular Radio morphology
-> 94 Candidates left
Select high radio luminosity and low luminosity at high frequencies
-> No candidates left

Outlook:

Include AGN from larger distances
Refine selection by:

- 1) More quantitative morphology selection
- 2) More detailed inspection of emission spectrum
 - 3) Develop new AGN selection criteria





Event selection + AGN selection There is more:

- Combined point-source analysis with the Auger and TA collaborations



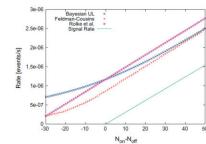




- Developing new statistical methods: Astropart. Phys. 50-52 (2013) 57-64 (ArXiv:1212.2008)

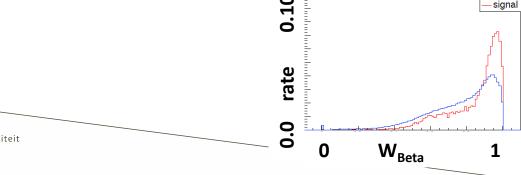
Dwalk





-burn

- Developing and implementing new IceCube track reconstruction algorithms





Summary

- So far IceCube did not see point sources: Strong constraints have been put on existing GRB models!
- At the Brussels IIHE AGN/GRB group a new IceCube event selection method is under development showing very promising initial results.
- A specific type of AGN (obscured Blazar) will be considered for the Brussels AGN analysis. A selection method for this specific type of AGN is currently under investigation
- Next to these subjects, the Brussels IIHE AGN/GRB group is also
 - involved in:
- <u>Combined point-source analysis with the Auger and TA</u> collaborations
- <u>Developing new statistical methods:</u>
 Astropart.Phys. 50-52 (2013) 57-64 (ArXiv:1212.2008)
- <u>Developing and implementing new IceCube track</u> reconstruction algorithms

